HRBE Data Analysis

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Outline

- Review/Background
- Magnetic Activity in the Data
- Comparing HRBE Flux and NOAA POES Flux
- Future Work

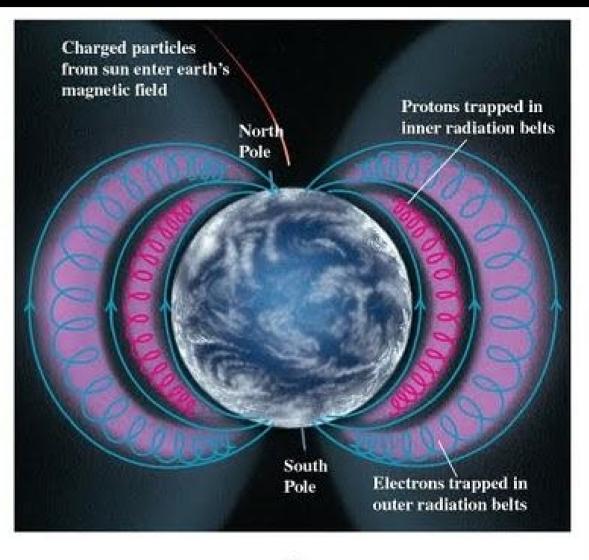
HRBE (aka Explorer -1 [Prime]



- Designed to monitor the intensity and location of the radiation belts
 - Flies a Geiger tube donated by Dr. James Van Allen
 - Counts charged particles, mainly electrons > 50 keV

Van Allen Radiation Belts

- 2 regions of trapped charged particles
- Located at ~1.5 R_{F} and ~4-5 R_{F}^{1}
- Dynamic and highly variable depending on solar activity
- Filled with electrons and protons
 - Protons significant only in inner belt
- Slot region (for Dana):
 - Inward radial diffusion and pitch angle scattering during quiet time
 - Whistler Mode Waves cause pitch angle scattering of trapped particles after storms
 - Particles lost to atmosphere



(a)

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Motion of Charged Particles²

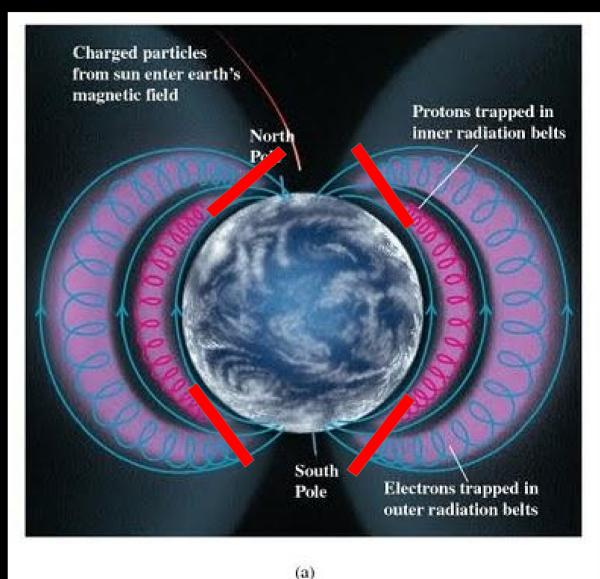
 $F = q (v \times B + E)$ **Bounce Motion:** Gyro Motion: The Magnetic Mirror Effect $m\iota$ \mathbf{F}_m В Gradient Drift mv v_G $2ab^3$ В **Curvature Drift** mv v_C $qB^{\overline{3}}$ **Drift Motion Bounce Motion** Gyro Motion

L-Shells

- L-Shell Coordinate system²
 - Eliminates longitudinal coordinate
 - Particle drifting and bouncing around earth has nearly L=constant

$$r = L\cos^2 \lambda$$
$$B = \frac{B_0}{r^3} (4 - \frac{3r}{L})^{\frac{1}{2}}$$

- r is in Earth radii, lambda is geomagnetic latitude
- L R_E ~ distance from dipole center to a field line at the equator

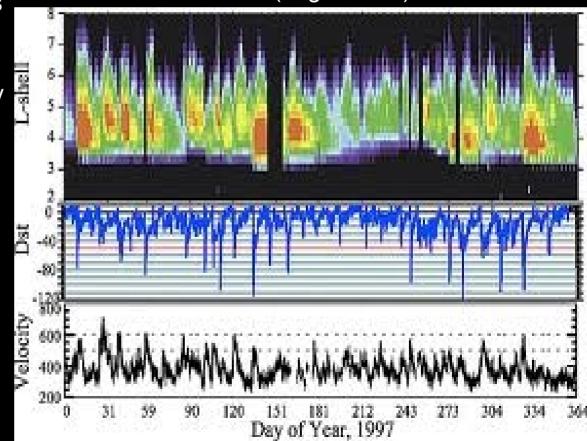


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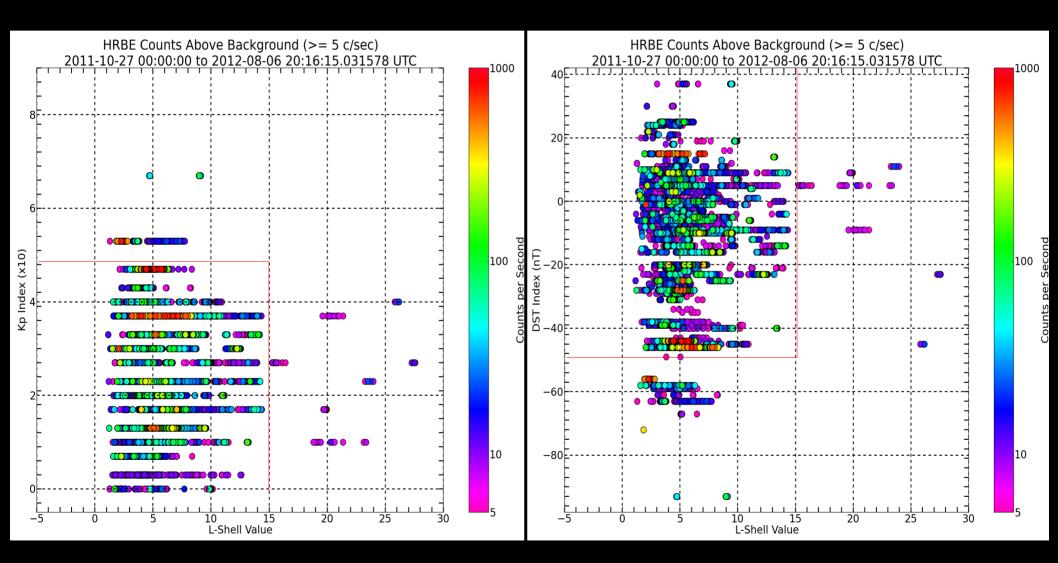
Kp and Dst Indexes

- Measures of the horizontal component of the earth's magnetic field
- Magnetic field varies with solar activity
- KP Index ³
 - Measured from higher latitudes
 - KP of 0-4 is below storm level
 - Kp of 5-9 indicates increasingly powerful storms

- DST Index ⁴
 - Measured close to the equator
 - DST < -50 nT indicates a magnetic storm
 - Measure of magnetospheric current (ring current)

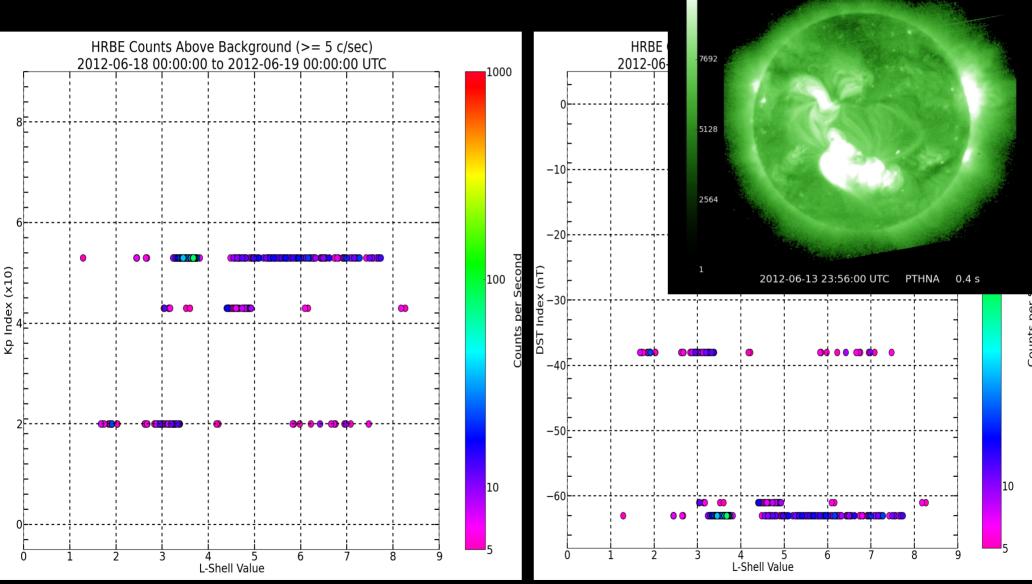


Geomagnetic Activity in HRBE Data

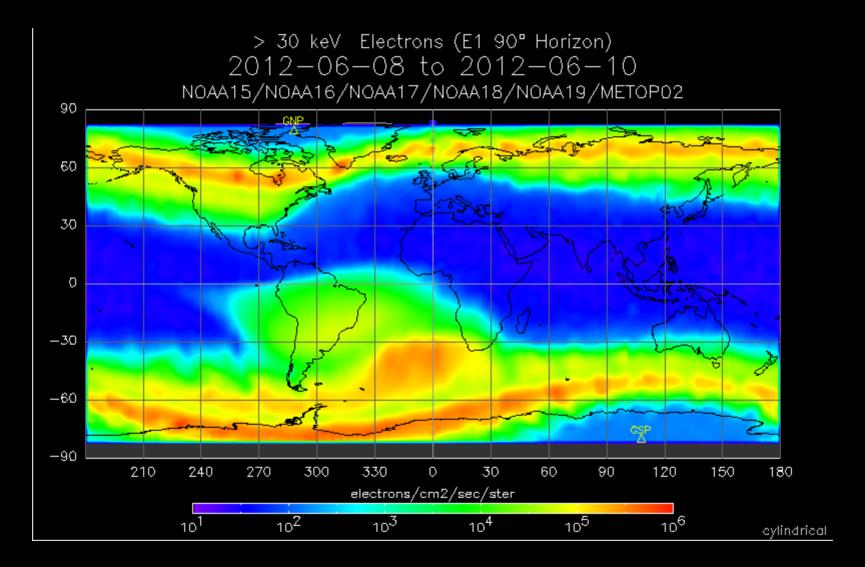


18 June: M-Class Flare

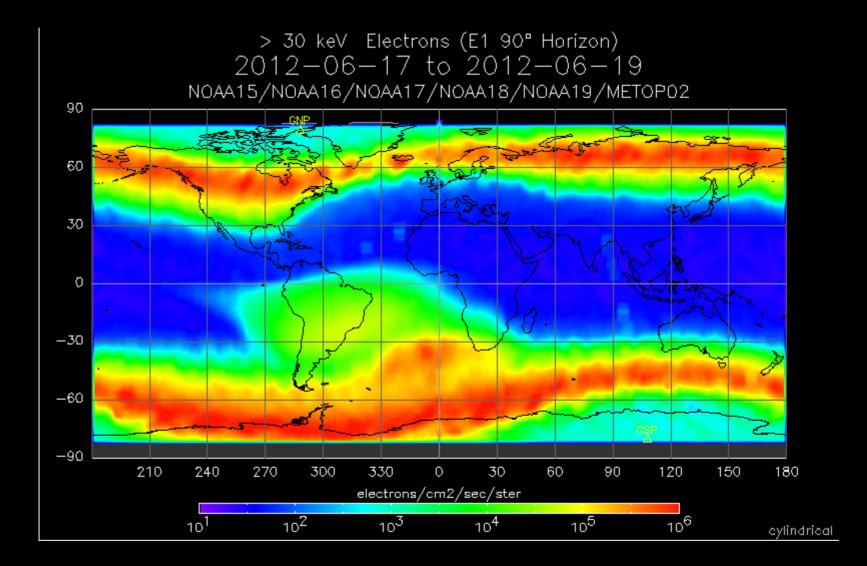
GOES-15 SXI Level-1 NOAA/SWPC Boulder, CO



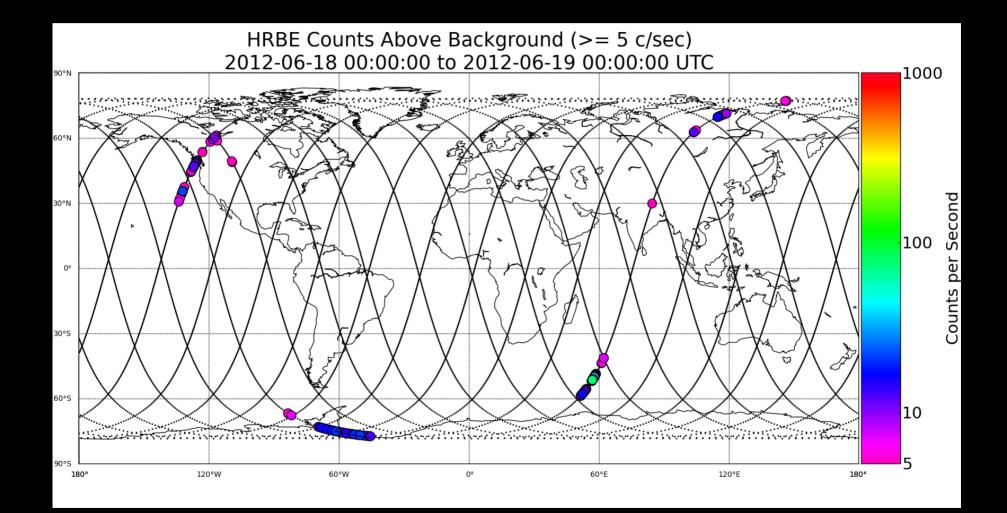
9 June: A Quiet Day



18 June: M-Class Flare



18 June: M-Class Flare



HRBE Flux

- HRBE software stores a count rate: 'Counts per second'
- Flux can be calculated:

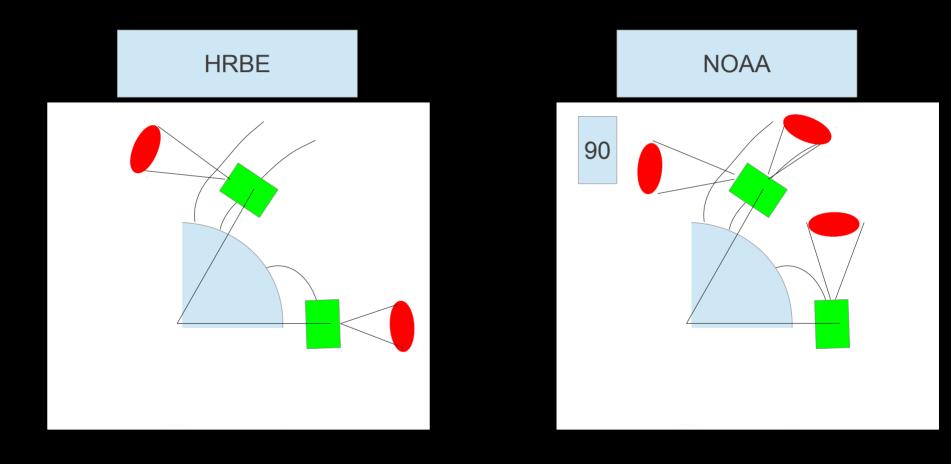
$$\Phi = \frac{Counts}{s} \frac{1}{GF \times Efficiency}$$

• GF is Geometric Factor:

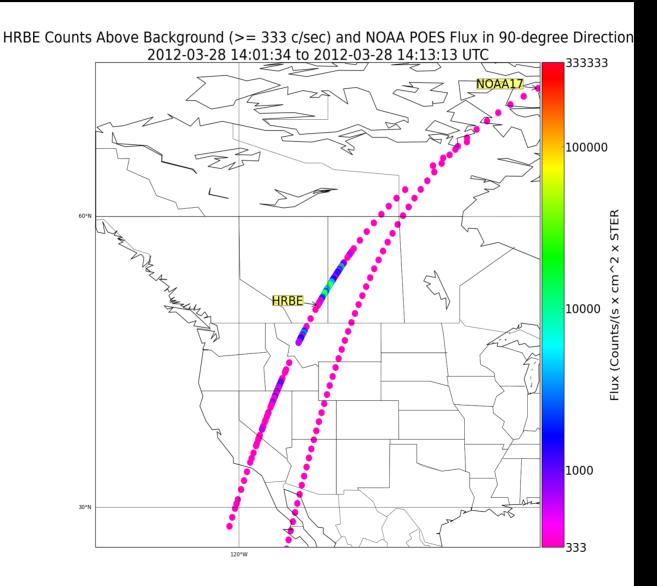
$$GF = Area \times \Omega$$

$$\Omega = SolidAngle$$

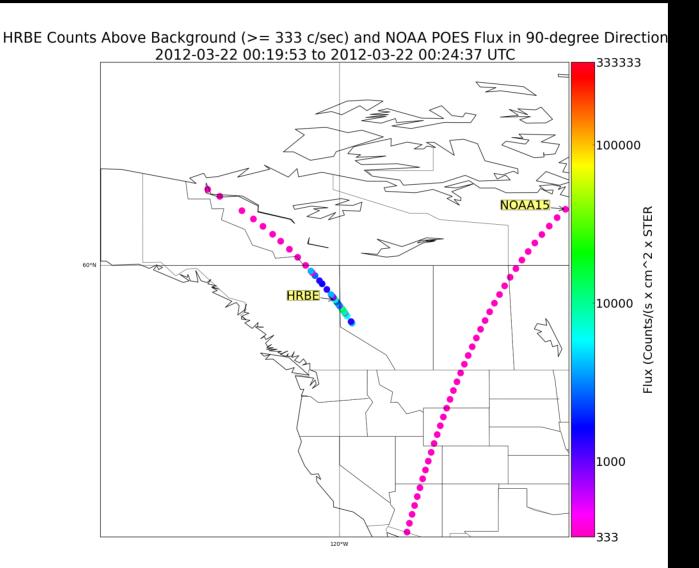
Why should HRBE be coincident with the 90-degree detector?



Direct Flux Comparisons: HRBE with NOAA POES Satellites



Direct Flux Comparisons: HRBE with NOAA POES Satellites



Further Work

- Compare our count data to SAPS events
- Continue comparing our flux with NOAA flux
 - Gain a better understanding of how the satellite is working
 - Attempt to normalize our flux with NOAA POES flux
 - Explain differences in our flux and NOAA flux
- Look for solar proton events appearing in our data
 - Data taken at high L-values (over open field lines)
- Write a paper

Citations

- 1. Li, X, and M.A. Temerin, The Electron Radiation Belt, *Space Science Reviews*, 95(1/2), 569, 2001.
- 2. Walt, Martin, Introduction to Geomagnetically Charged Particles, Cambridge University Press, 2005.
- 3. http://www.swpc.noaa.gov/info/Kindex.html
- 4. Temerin, M.A., and X. Li, A New Model for the Prediction of Dst on the Basis of the Solar Wind, *Journal of Geophysical Research*, Vol 107 (A12), 1472, 2002.

Acknowledgements

- Special thanks to Dave for all his help and guidance
- Matt Handley for answering all my computer questions
- The guys in SSEL
- The REU crew