



FIRS: A New Instrument for Photospheric and Chromospheric Studies at the Dunn Solar Telescope

Facility InfraRed Spectropolarimeter

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This project is supported by the National Science Foundation Major Research Program, Award No. ATM-0421582.





S.Jaeggli for NSO Workshop 25

Solar B, Photosphere to Corona

- Coronal heating
- Vertical wave propagation from the photosphere to chromosphere
- Running penumbral waves
- Normal and inverse Evershed flows
- 3D magnetic structure of sunspots
- Prominence and spicule magnetic fields
- Canopy fields
- Dynamic events and evolution of everything with time







Multi-Spectral: Multi-Height Diagnostics

Line [Å]	Species	Landé g	Formation Height
6302	Fe I	1.667, 2.5	mid-photosphere
8542	Ca II	1.1	low chromosphere
10827	Si I	1.5	photosphere
10830	He I	2.0, 1.75, 1.25	high chromosphere
15650	Fe I	3, 1.53	low photosphere





Multi-Spectral: Multi-Height Diagnostics

High Cadence: Dynamics





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High Cadence: Dynamics

High Spectral Resolution: Detailed Physics





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High Cadence: Dynamics

High Spectral Resolution: Detailed Physics

High Spatial Resolution: Structure Detail





Instrument Requirements with FIRS





Multi-Spectral: Multi-Height Diagnostics

Dual-arm spectrograph for simultaneous visible and infrared observations

Beam splitter to share the beam with IBIS

		Line [Å]	Species	Landé g	Formation Height
	*	* 6302	Fe I	1.667, 2.5	mid-photosphere
with IBIS	*	* *8542	Ca II	1.1	low chromosphere
k	*	10827	Si I	1.5	photosphere
	*	10830	He I	2.0, 1.75, 1.25	upper chromosphere/low corona1500-2000 km
		<mark>*</mark> 15650	Fe l	3, 1.53	low photosphere





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Building a Scan with 4 Slits







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High Cadence: Dynamics

Increased throughput with 4 slits

High Spectral Resolution: Detailed Physics

Large IR grating with steep blaze 300,000 for the IR, 600,000 for the visible (3pm measured with HeNe)





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High Spatial Resolution: Structure Detail

Diffraction-limited with HOAO

f/36 and f/108 feed optics for high and low-res modes









FIRS Properties

* Assuming use of 40 μm slit ** S/N ~ 10³

Property	FIRS f/36*	FIRS f/108*	Hinode SOT/SP
Telescope	76.2 cm Solar Tower		50 cm Aplanatic Gregorian
Rayleigh limit @ 6302	0.21"	•••	0.32"
Rayleigh limit @ 10830	0.36"		
Rayleigh limit @ 15648	0.52"		•••
Field	174" x 75"	58" × 25"	160" (320" max) x 151"
Vis Spatial Sampling	0.30" x 0.08"/pix	0.10" x 0.03"/pix	0.15" x 0.16"/pix
IR Spatial Sampling	0.30" x 0.15"/pix	0.10" x 0.05"/pix	
Nominal Scan Time**	20 min		83 min
6302 Spectral Resolution (Sampling)	0.03 (0.01) Å		0.03 (0.02) Å
10830 Spectral Resolution (Sampling)	(0.04) Å		
15648 Spectral Resolution (Sampling)	0.17 (0.05) Å		

FIRS Data





First Light 30 April, 2007

So far: 6 regions, 20 days, various configurations

Best yet: NOAA 11024, 7 July, 2009

IRS



2009-07-07 14:31:00 UT



2009-07-07 18:02:00 UT





[arcsec]

20

Hinode scan started 2009-07-07 14:03:00 UT



Current Status

FIRS standard configuration is complete!

- Observer training July, September 2009
- Released for NSO use 4th quarter 2009
- Released for general use 1st quarter 2010





For the Future

Extended Capabilities:

- Additional visible and IR detectors for simultaneous 6302-8542-10830-15650
- Maybe super-achromatic dual-waveplate modulator for synchronized exposures
- New pair of Wollaston prisms for smaller beam deviation and larger FOV
- More narrow-band filters for extended wavelength coverage





Thank You!

Ask us for a tour

for more information visit: http://kopiko.ifa.hawaii.edu/firs/

> or contact us: jaeggli@ifa.hawaii.edu lin@ifa.hawaii.edu





