# Solar Coronal Tomography

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#### **Abstract**



A simple, yet powerful, algorithm for computed tomography of the solar corona is demonstrated using synthetic EUV data. A minimum of three perspectives are required. These may be obtained from STEREO/EUVI plus an instrument near Earth, e.g. TRACE or SOHO/EIT.



# Coronal Tomography (EUV)

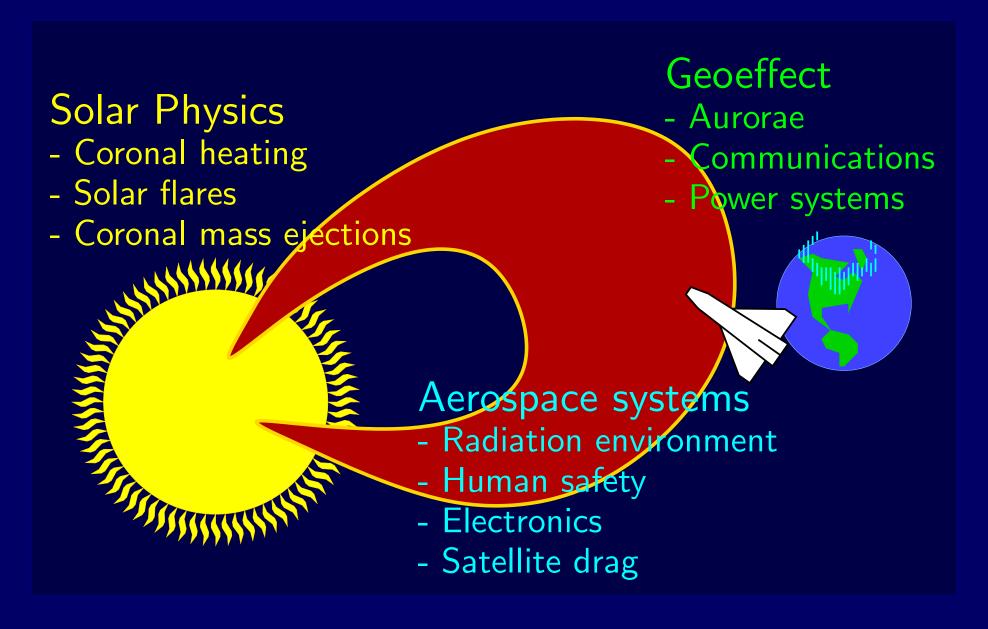


- 1. Why is the Solar Corona important?
- 2. Background—STEREO mission
- 3. Science Goals
- 4. SMART Algorithm
- 5. Demonstration (synthetic EUV data)
- 6. Summary & conclusions



# Why is the Corona interesting?



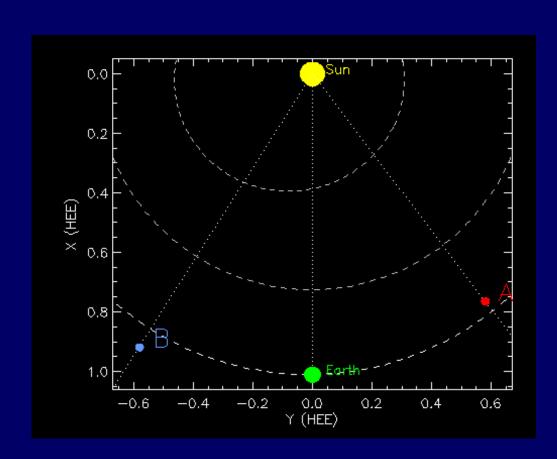




#### NASA's STEREO Mission



(SECCHI remote sensing pkg)



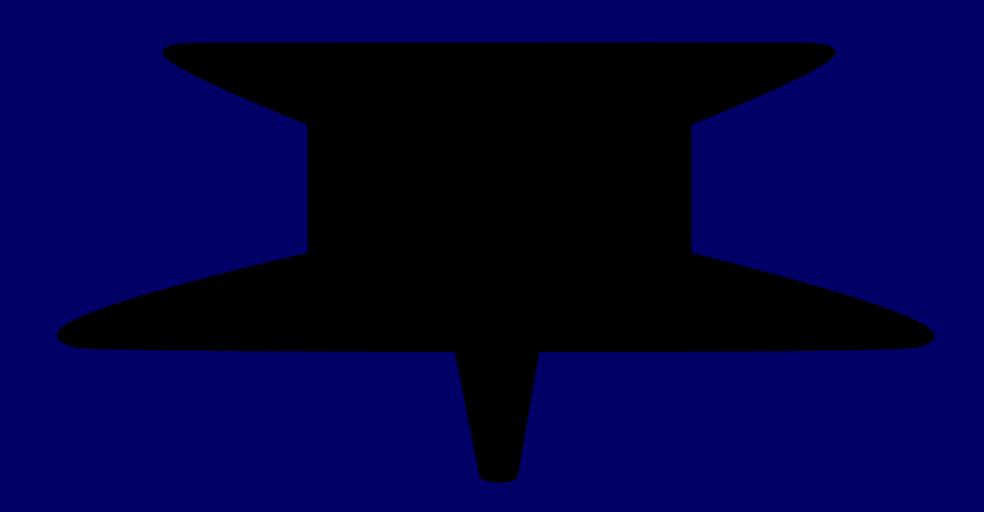
 $STEREO \text{ Aug 27 } (-32^{\circ}, +37^{\circ})$ 

- Twin spacecraft in solar orbit
- $\bullet$  A = Ahead, B = Behind
- EUVI = Extreme UV Imager (low corona)
- Coronagraphs observe outer corona
- Heliospheric Imager
- Like human stereoscopic vision (?)



# Example EUVI Data (171 Å)







## Science Objectives



A primary science question of STEREO/SECCHI is:

What is the 3-dimensional structure of active regions, coronal loops, helmet streamers, etc?<sup>1</sup>

#### Approach:

- $\bullet$  Adapt Smooth Multiplicative Algebraic Reconstruction Technique (SMART) algorithm developed for MOSES.
- 3 views: STEREO/EUVI + TRACE (EUV telescope in LEO, launched April 1998).

<sup>&</sup>lt;sup>1</sup>Howard et al. 2007, Space Science Reviews



### Inspiration from MOSES



What does this sounding rocket have to do with STEREO?

- Multi-Order Solar EUV Spectrograph (MOSES)
- NASA LCAS launch, 2006 Feb 8
- Simultaneous imaging and spectroscopy
- Tomographic inversion → spectrum
- ullet Synergy with STEREO few-angle tomography





#### **Problem Statement**



Observe volume EM I(x,y,z) at angles  $\theta_i$  by projection:

$$f_i(x, z) = \int_D \mathcal{R}_z(\theta_i) I \, dy + \text{noise}$$

Given images  $f_i$  and initial guess G(x, y, z), iteratively modify G to approximate I.









#### 1. Smooth G.





- 1. Smooth G.
- 2. Project:  $f'_i = \int_D \mathcal{R}_z(\theta_i) G dy \quad \forall i$ 
  - ullet Evaluate goodness of fit,  $\chi^2_{R,i}$
  - ullet Adjust  $\gamma_i$  based on  $\chi^2_{R,i}$  and  $\Delta\chi^2_{R,i}$





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- 3. Correct:  $G \leftarrow G \prod_i C_i^{\gamma_i}$ , where

$$C_i(x,y,z) = \mathcal{R}_z(- heta_i) rac{f_i(x,z)}{f_i'(x,z)}.$$





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$$C_i(x,y,z) = \mathcal{R}_z(-\theta_i) \frac{f_i(x,z)}{f'_i(x,z)}.$$

Iterate, adjusting  $\gamma_i$  to drive  $\chi^2_{R,i} \to 1$ .

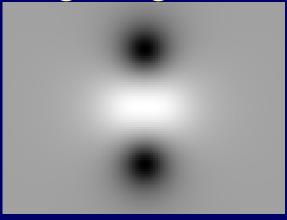


## Synthetic Data

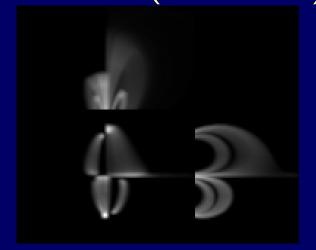


- Quadrupole potential field
- Coronal null: interesting geometry
- EM along arbitrary field lines
- Random latitude, longitude & clocking
- 3 viewpoints:  $-40^{\circ}, 0^{\circ}, +40^{\circ}$
- Imaged with 3000 counts in the brightest pixel
- Note square root scaling

#### Magnetogram:



#### Corona (isometric):





## Example 1 (E-W pan)





# Example 1 (N-S pan)





# Example 2 (E-W pan)





# Example 2 (N-S pan)





## **Summary & Conclusions**



- Full 3D reconstruction of coronal EM
- 3 viewpoints required
- No assumptions (loops, magnetic field, etc.)
- Simple and computationally efficient
  - Examples converged in 15 iterations
  - 3 minutes on my laptop ( $169^3$  voxels)
- Loops parallel to ecliptic—not so good
- Apply to EUVI data hopefully this fall!

Acknowledgment: This work has been funded through the Solar and Heliospheric LCAS Program.