Space Weather Highlights

20 - 26 October 1997

Solar activity was very low to low. Isolated B-class subflares occurred during most of the period. A 22 degree-long filament disappeared from the southeast quadrant late on 20 October. A long-duration B-class X-ray flare and a weak Type II radio sweep were associated with the disappearance. A space-based sensor also detected a partial-halo coronal mass ejection (CME) coincident with the disappearance. Activity rose to low levels on 21 October with the occurrence of a C3/SF flare at 21/1754UT. A halo CME was associated with this flare. Activity declined to very low levels after 21 October.

Solar wind data were received from the WIND spacecraft a few hours per day. Data were unreliable during 21 - 22 October due to spacecraft location within the magnetopause. Velocities were steady at around 300 km/sec early in the period, then gradually increased to 550 km/sec during the rest of the period. Solar wind particle densities were enhanced during 23 - 24 October peaking near 60 p/cc on the 24th. Densities dropped to around 05 p/cc during 25 - 26 October. Bz was in the plus to minus 10 nT range (GSM) during most of the period, but varied from minus 23 to plus 17 nT during 24 October. Solar sector orientation was mostly away (phi angle near 135 degrees) on 20 October, but was not discernible during 23 - 24 October. Orientation was mostly away during 25 - 26 October.

No significant proton enhancements were observed at geosynchronous altitude.

The greater than 2 MeV electron flux at geosynchronous altitude reached high levels late on 26 October.

The geomagnetic field was quiet to unsettled through late 23 October. Activity increased after 23/2100UT with active to minor storm levels observed at all latitudes and isolated major storm periods at high latitudes. The increased activity persisted through the remainder of the interval.

Space Weather Forecast 29 October 1997 - 24 November 1997

Solar activity is expected to range from very low to low.

No significant proton enhancements are expected at satellite altitudes.

The greater than 2 MeV electron flux at geosynchronous altitude is expected to be normal to moderate during most of the period.

The geomagnetic field is expected to be mostly quiet to unsettled.



Daily Solar Data

	Radio	Sun	Sunspot	X-ray				Flares				
	Flux	spot	Area	Background	X	ray Fl	ux		Op	tical		
Date	10.7 cm	No. (2	10 ⁻⁶ hemi.)		C	M	X	S	1	2	3	4
20 October	83	25	60	A8.1	0	0	0	0	0	0	0	0
21 October	85	27	50	A5.8	1	0	0	1	0	0	0	0
22 October	81	13	10	A7.2	0	0	0	0	0	0	0	0
23 October	80	0	0	A2.4	0	0	0	0	0	0	0	0
24 October	79	11	0	A2.0	0	0	0	0	0	0	0	0
25 October	81	13	20	A2.3	0	0	0	0	0	0	0	0
26 October	82	14	30	A5.5	0	0	0	0	0	0	0	0

Daily Particle Data

			Dany Far	ucie Daia						
		Proton Fluence otons/cm ² -day			Electron Fluence (electrons/cm²-day-sr)					
Date	>1MeV	>10MeV	>100MeV	>.6MeV	>2MeV	>4MeV				
20 October	6.2E + 5	1.9E + 4	5.0E+3		1.7E+6					
21 October	1.1E+6	1.9E + 4	5.2E+3		1.9E+6					
22 October	6.7E + 5	2.1E+4	4.8E+3		4.2E + 5					
23 October	7.5E + 5	1.9E + 4	4.6E + 3		2.9E + 5					
24 October	1.6E + 5	1.7E + 4	4.0E + 3		1.2E + 5					
25 October	3.8E + 5	1.6E+4	3.6E+3		2.9E+6					
26 October	6.1E+5	1.6E+4	3.5E+3		4.6E+7					

Daily Geomagnetic Data

	M	Iiddle Latitude	I	High Latitude		Estimated
	F	Fredericksburg		College		Planetary
Date	A	K-indices	A	K-indices	A	K-indices
20 October	5	2-1-3-1-1-1-1	2	0-0-3-0-1-0-0-0	6	2-1-3-1-1-2-1-1
21 October	3	1-1-1-1-1-1	2	0-0-2-3-0-0-0	4	1-0-1-2-2-2-1-1
22 October	7	3-3-1-1-1-2-2	0	0-0-0-0-0-0-0	5	2-2-1-1-2-2-2
23 October	6	0-1-2-2-1-2-2-3	3	0-0-1-0-0-1-1-3	6	0-0-1-2-2-2-3
24 October	18	2-3-3-4-4-3-2-4	14	1-2-4-5-3-1-1-2	25	2-3-3-5-6-3-2-3
25 October	14	4-4-2-3-2-2-3	14	3-3-2-4-4-3-1-1	19	5-5-2-4-3-3-2-3
26 October	15	4-3-1-1-3-3-3-4	12	3-1-0-0-4-5-2-1	14	4-3-0-1-3-4-3-4

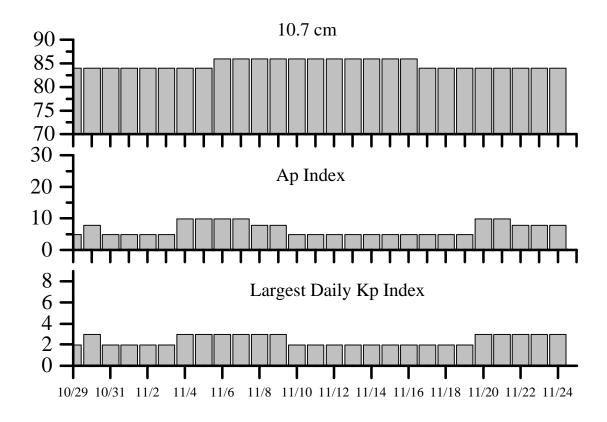


Alerts and Warnings Issued

Date and Time of Issue (UT)	Type of Alert or Warning	Date and Time of Event (UT)
20 Oct 0904	K=4 observed	20 Oct 06-09
21 Oct 1446	Type II Radio Emission	21 Oct 0107
22 Oct 1650	$A \ge 20$ Watch	25 Oct
23 Oct 0903	Sudden Impulse observed at Boulde	r 23 Oct 0830
23 Oct 1813	$A \ge 30$ Watch	26 Oct
24 Oct 1512	K = 4 Observed	24 Oct 12-15
24 Oct 1740	$A \ge 20$ Watch	24 Oct
25 Oct 0338	K= 4 Observed	25 Oct 00-03
25 Oct 0605	K= 5 Observed	25 Oct 03-06
25 Oct 0607	$A \ge 20$ Observed	25 Oct 0600
26 Oct 0304	K= 4 Observed	26 Oct 00-03
26 Oct 1637	> 2MeV Electron Event ≥ 1000pfu	26 Oct 1628



Twenty-seven Day Outlook



	Radio Flux	Planetary	Largest		Radio Flux	Planetary	Largest
Date	10.7 cm	A Index	Kp Index	Date	10.7 cm	A Index	Kp Index
29 Oct	84	5	2	11 Nov	86	5	2
30	84	8	3	12	86	5	2
31	84	5	2	13	86	5	2
01 Nov	84	5	2	14	86	5	2
02	84	5	2	15	86	5	2
03	84	5	2	16	86	5	2
04	84	10	3	17	84	5	2
05	84	10	3	18	84	5	2
06	86	10	3	19	84	5	2
07	86	10	3	20	84	10	3
08	86	8	3	21	84	10	3
09	86	8	3	22	84	8	3
10	86	5	2	23	84	8	3
				24	84	8	3



Energetic Event

	Time (UT)	X-ray	Optical Inform	nation	Peak	Sweep Freq
Date	1/2	Integ	Imp Location	Rgn	Radio Flux	Intensity
	Begin Max Max	Class Flux	Brtns Lat CMD	#	245 2695	II IV

No Event Observed

Flare List

				Flare List			
				_	О	ptical	
		Time		X-ray	Imp /	Location	Rgn
Date	Begin	Max	End	Class.	Brtns	Lat CMD	#
20 October	0219	0222	0226	B1.2			
	0330	0335	0338	B5.1			
	1110	1117	1132	B1.7			
	1213	1216	1219	B1.2			
	1350	1355	1402	B1.9			
21 October	0045	0236	0247	B3.4			
	1733	1750	1819	C3.3	SF	N16E07	8097
22 October	No Fla	res Obser	ved				
23 October	No Fla	res Obser	ved				
24 October	No Fla	res Obser	ved				
25 October	2035	2041	2058	B1.7			
26 October	0221	0227	0232	B1.3			
	2245	2249	2251	B1.9			
•							

Region Summary

	Location		Sunspot	Characte	ristics			F	lares			
	Helio	Area	Extent	Spot	Spot	Mag	X-ray		(Optic	al	
Date	(°Lat°CMD) Lon	(10 ⁻⁶ hemi)	(helio)	Class	Count	Class	C M X	S	1	2	3	4

Region 8095

12 Oct N33E42 215 0010 03 BXO 002 B 13 Oct N33E32 212 0000 01 AXX 002 A

14 Oct N33E19 212

15 Oct N33E06 212

16 Oct N33W07 212

17 Oct N33W20 212

0 0 0 0 0 0 0 0

Died on Disk.

Absolute heliographic longitude: 212



	Location	1		Sunspot (Sunspot Characteristics						Flares				_
D	0.T 0.GP. (TD.)	Helio	Area	Extent	Spot	Spot	Mag		X-ray		_		ptica		_
Date (° Lat ° CMD) Region	Lon n 8096	(10 ⁻⁶ hemi) (helio)	Class	Count	Class	<u>C</u>	M	X	S	1	2	3	<u>4</u>
14 Oct	N14E62	169	0020	04	BXO	003	В				1				
	N15E47	171	0050	06	CSO	006	В								
	N14E34	171	0030	06	BXO	005	В								
	N13E21	171	0040	07	BXO	010	В								
	N13E08	170	0030	06	CSO	007	В								
	N13W03	167	0030	04	CSO	004	В								
	N13W16	167	0020	02	HSX	004	Ā								
	N13W29	167	0010	03	BXO	003	В								
	N13W43	168	0000	02			_								
	N13W56	168													
	N13W69	168													
	N13W82	168													
								0	0	0	1	0	0	0	0
Crossec	l West Lim	b.													
Absolut	te heliograp	hic long	itude: 10	57											
	<i>C</i> 1		,												
	Region	ı	8097												
15 Oct	N18E77	141	0060	02	HSX	002	A								
16 Oct	N16E69	136	0060	08	CSO	002	В				1				
17 Oct	N17E57	135	0080	08	CSO	003	В								
18 Oct	N17E43	135	0060	08	CSO	003	В								
19 Oct	N17E26	138	0030	02	HSX	001	A								
20 Oct	N16E13	138	0040	02	HSX	001	A								
21 Oct	N16W01	139	0040	03	CRO	004	В	1			1				
22 Oct	N16W14	139	0010	02	BXO	003	В								
23 Oct	N16W27	139													
24 Oct	N16W40	139													
25 Oct	N16W53	139													
26 Oct	N16W66	139													
								1	0	0	2	0	0	0	0
Still on															
Absolut	te heliograp	_		39											
		gion 80:													
	S26W03	101	0000	00	AXX	001	A								
	S26W16	101													
26 Oct	S26W29	101						_	-	_	_	_	_	-	
a	D. 1							0	0	0	0	0	0	0	0
Still on			• •	24											
Absolut	te heliograp	hic long	gitude: 10	01											



Region Summary-continued.

	Extent (helio)	Spot Class	Spot Count	Mag Class	C	X-ray M	X	S	1	ptica 2	al 3	4
8099	(helio)	Class	Count	Class	С	M	X	S	1	2	3	4
												<u> </u>
1 0000		Region 8099										
1 0020	04	CRO	003	В								
0030	06	CSO	004	В								
					0	0	0	0	0	0	0	0
Absolute heliographic longitude: 020												
						0	0 0	0 0 0	0 0 0 0	0 0 0 0 0	0 0 0 0 0 0	0 0 0 0 0 0 0



Recent Solar Indices (preliminary) of the observed monthly mean values

					monthly .	mean value				
			ot Numbers				lio Flux	Geomagnetic		
-	Observed		Ratio	Smooth		Penticton	Smooth	-		
Month	SWO	RI	RI/SWO	SWO	RI	10.7 cm	Value	Ap	Value	
					400=					
					1995					
October	31.6	21.1	0.67	19.7	12.1	77.9	73.8	16	11.5	
November	15.7	09.0	0.57	18.5	11.4	74.2	73.2	09	10.8	
December	16.2	10.0	0.62	17.6	10.8	72.6	72.8	09	10.0	
					1996					
January	17.6	11.5	0.65	16.8	10.4	74.5	72.4	09	09.8	
February	09.1	04.4	0.48	16.2	10.1	71.5	72.2	10	09.8	
March	12.1	09.2	0.76	15.4	09.7	70.7	72.1	11	09.9	
April	08.5	04.8	0.56	13.6	08.5	69.3	71.6	11	09.7	
May	11.8	05.5	0.47	12.9	08.0	70.1	71.4	07	09.5	
June	18.8	11.8	0.63	13.5	08.5	69.6	71.8	05	09.4	
July	13.2	08.2	0.62	13.4	08.4*	71.2	72.0	07	09.3	
August	20.5	14.4	0.70	13.1	08.3*	72.4	72.1	09	09.4	
September	02.9	01.6	0.55	13.3	08.5*	69.4	72.3	15	09.3	
October	02.3	00.9	0.39	14.0	08.9*	69.2	72.6	13	09.1	
November	26.7	17.9	0.59	15.4	09.9*	78.7	73.0	08	09.1	
December	21.1	13.3	0.63	16.2	10.5*	78.7 77.8	73.0	08	09.1	
December	21.1	13.3	0.03		10.5 · 1997	11.0	13.3	07	09.3	
January	09.0	06.5*	0.72*	16.5	10.6*	74.0	73.4	09	09.3*	
February	11.3	07.6*	0.67*	17.4	11.2*	73.8	73.7	11	09.2*	
March	14.4	08.8*	0.61*	20.4	13.7*	73.5	75.1*	08	09.0*	
April	24.5	15.8*	0.64*			74.5		10		
May	28.6		0.65*			74.6		08		
June	22.1	13.1*	0.59*			71.7		07		
July	17.0	10.5*	0.62*			71.1		06*		
August	36.7	24.7*	0.67*			79.0		08*		
September	58.2	51.3*	0.88*			96.2*		10*		
1			-							

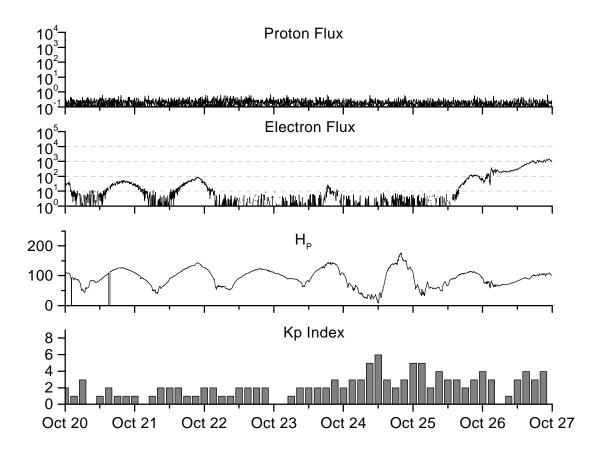
^{*}Preliminary estimates.

October 1996 has been selected as the start of Solar Cycle 23.

The highest smoothed sunspot number for Solar Cycle 22, RI=158.5, occurred in July 1989.

The highest smoothed 10.7cm Radio Flux value for Solar Cycle 22, of 213.1, occurred in June 1989.





Weekly Geosynchronous Satellite Environment Summary Week Beginning 20 October 1997

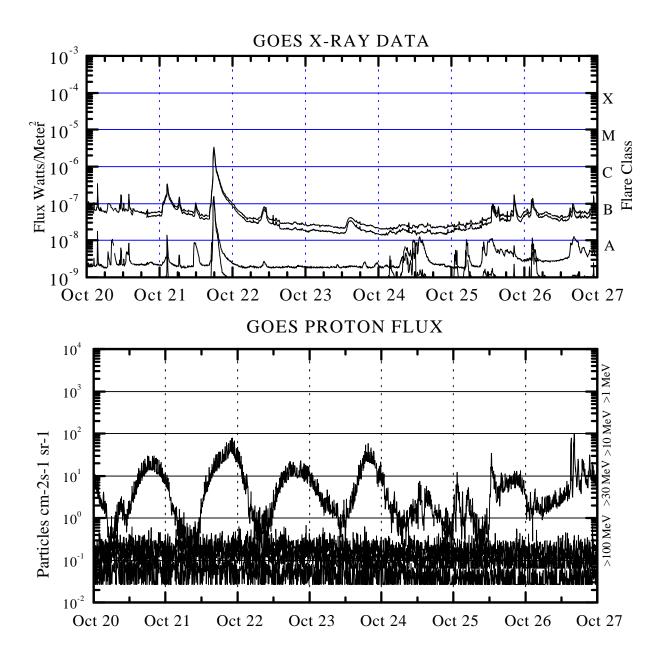
Protons plot contains the five-minute averaged integral proton flux (protons/ cm²-sec-sr) as measured by GOES-9 (W135) for each of three energy thresholds: greater than 10, 50, and 100 MeV.

Electrons plot contains the five-minute averaged integral electron flux (electrons/ cm^2 -sec-sr) with energies greater than 2 MeV at GOES-9.

Hp plot contains the five minute averaged magnetic field H component in nanoteslas (nT) as measured by GOES-9. The H component is parallel to the spin axis of the satellite, which is nearly parallel to the Earth's rotation axis. *Kp* plot contains the estimated planetary 3-hour K-index (derived by the USAF 55th Space Weather Squadron) in real time from magnetometers at Meanook, Canada; Sitka, AK; Glenlea, Canada; St. Johns, Canada; Ottawa, Canada; Newport, WA; Fredericksburg, VA; Boulder, CO; Fresno, CA and Heartland, UK. These data are made available through cooperation from the Geological Survey of Canada (GSC) and the US Geological Survey. These may differ from the final Kp values derived from a more extensive network of magnetometers.

The data included here are those now available in real time at the SWO and are incomplete in that they do not include the full set of parameters and energy ranges known to cause satellite operating anomalies. The proton and electron fluxes and Kp are "global" parameters that are applicable to a first order approximation over large areas. Hparallel is subject to a more localized phenomena and the measurements generally are applicable to within a few degrees of longitude of the measuring satellite.



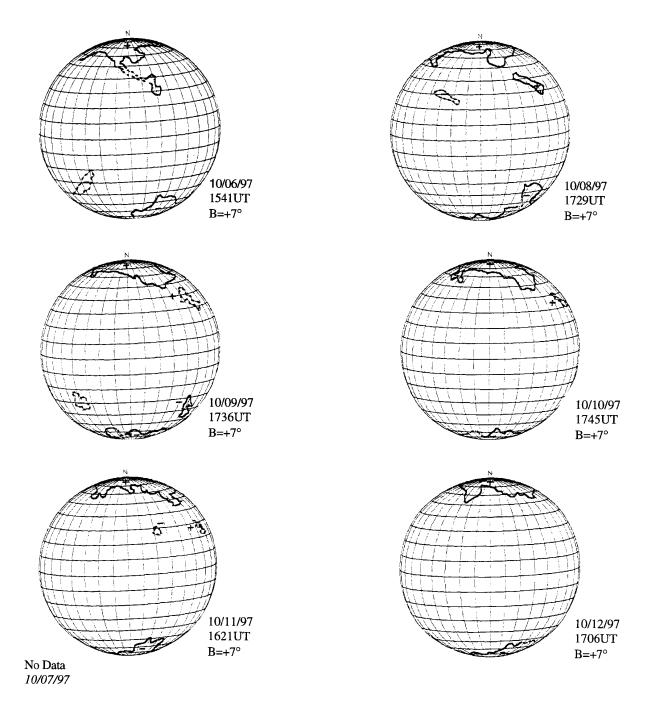


Weekly GOES Satellite X-ray and Proton Plots

Proton plot contains the five minute averaged integral proton flux (protons/cm 2 -sec-sr) as measured by GOES-9 (W135) for each of the energy thresholds: >1, >10, >30 and >100 MeV. P10 event threshold is 10 pfu (protons/cm 2 -sec-sr) at greater than 10 MeV.

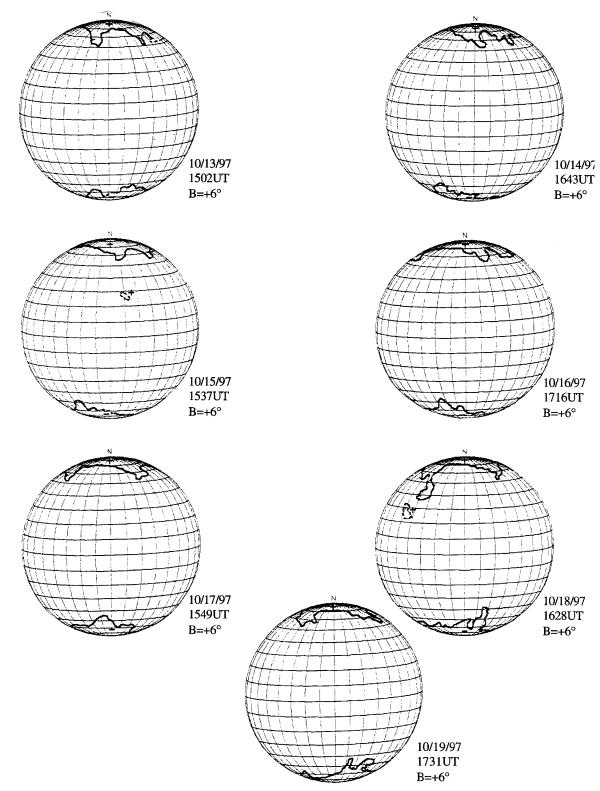
X-ray plot contains five minute averaged x-ray flux (watts/m²) as measured by GOES 8 and 9 in two wavelength bands, .05 -.4 and .1 - .8 nm. The letters A, B, C, M and X refer to x-ray event levels for the .1 - .8 nm band.





Coronal hole maps from the National Solar Observatory, Kitt Peak, Arizona
These maps are reproductions of the coronal hole contours as derived from the Kitt Peak 1083 nm raw data images. These are preliminary data, where solid (and dashed) lines are used to indicate more (and less) confidence in an inference and are printed whenever they are available at SWO.





Coronal hole maps from the National Solar Observatory, Kitt Peak, Arizona
These maps are reproductions of the coronal hole contours as derived from the Kitt Peak 1083 nm raw data images. These are preliminary data, where solid (and dashed) lines are used to indicate more (and less) confidence in an inference and are printed whenever they are available at SWO.

