

Solar Influence on climate:

Particle precipitation effects on the southern hemisphere tropical/subtropical lower stratosphere temperature

Luis Eduardo Vieira and Ligia Alves da Silva



Outline

- Overview
- Southern Hemisphere Lower stratosphere temperature distribution
- Vertical distribution of O3 and NO inside and outside the magnetic anomaly
- Southern Hemisphere climate sensitivity to zonal ozone asymmetry
- Conclusions

Sun-Climate Mechanisms

- The variability of the total solar irradiance causing a change in the total energy input to the earth's atmosphere and consequent warming/cooling;
- The variability of the solar ultraviolet emission and its effects on the stratospheric ozone and thermal structure;
- High energy particle precipitation effects on mesospheric and stratospheric ozone in the auroral and/or southern hemisphere magnetic anomaly regions; and,
- The cosmic rays effects on the cloud coverage.

Geomagnetic Modulation

Motivation: Increase of global surface temperature



Sea-level Pressure



 The ISCCP D2 data/images were obtained from the International Satellite Cloud Climatology Project web site http://isccp.giss.nasa.gov maintained by the ISCCP research group at the NASA Goddard Institute for Space Studies, New York, NY.



Source: Zhang, M., and H.

Song (2006), Evidence of

overturning circulation over the tropical Pacific,

Geophys. Res. Lett., 33,

doi:10.1029/2006GL0259

deceleration of atmospheric vertical

L12701,

42.

FD 1983-2004 Mean Annual

Long-standing debate in the climate community



How the **tropical Pacific** will respond to increase greenhouse gases:

Will the structure of changes in the ocean surface more closely resemble an El Nino or a La Nina?

(Vecchi et al., EOS, Transactions, v. 89, February 2008)

Sea-Surface Temperature TRMM Microwave Imager (**TMI**) satellite observations

The geographical distribution of the quasi trapped electron fluxes for an energy of 200 keV, as deduced from measurements by the IDP instrument



Long-wavelength cloud effects on the radiative flux in the atmosphere



•Annual means of the long-wavelength effects on radiative flux in the atmosphere for the period July 1983 to June 2001. The superposed black lines show the iso-intensity contours of the geomagnetic field at ground for year 1990.

•The ISCCP D2 data/images were obtained from the International Satellite Cloud Climatology Project web site http://isccp.giss.nasa.gov maintained by the ISCCP research group at the NASA Goddard Institute for Space Studies, New York, NY.

Vieira and Da Silva (2006), GRL

SHMA Evolution



IGRF MODEL Year: 2000







IGRF Model

The evolution of the Sea-Level Pressure in the Pacific SHMA region



Figure: It is shown in the upper panel the climatology of the sealevel pressure from 1948-2005. The superposed black lines show the difference of the geomagnetic field intensity at surface between years 2000 and 1940.

Figure: It is shown in the left upper panel the scatter plot of the Sea-level Pressure versus the Magnetic field intensity. The black line shows the best linear fit for the whole data set. The dotted black lines show the 1-sigma model error. It is shown in the right panel the frequency distribution of the Model Error.

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The Microwave Sounding Units (MSU) Data



The Microwave Sounding Units (MSU) Data

K Jun, 1979-2005 225 20° S 220 215 40° S 210 205 60 90° E 180[°] W 90° W 90° E 0° 200 κ Jul, 1979-2005 225 20[°] S 220 215 40[°]S 210 205 60 180° W 90° W 90° E O° 90° E 200 κ Aug, 1979-2005 225 220 20[°] S 215 40[°] S 210 205 60 90° E 180° W 90° W 90° E 0° 200 K Sep, 1979-2005 225 20[°] S 220 215 40[°] S 210 205 60 90° E 90° W 90° E 180° W 0° 200 Κ Oct, 1979-2005 225 20[°] S 220 215 40° S 210 205 60 90° E 90° E 180° W 90° W O° 200 Κ Nov, 1979-2005 225 220 20[°] S 215 40° S 210 205 60 90° E 180° W 90° W 90° E 0° 200

Southern hemisphere lower stratosphere temperature climatology for June to November. The superimposed black lines show the iso-intensity contours of the geomagnetic field at 10 km for year 1990.

Da Silva et al. (2008), GRL, submitted



Figure:

Longitudinal profiles of the lower stratosphere brightness temperature and the magnetic field intensity for November for the latitudes 80o S, 55o S, 42o S and 35o S.

Monthly observations are shown in light gray '+'. The continuous outsized lines show the mean temperature profiles and the tiny lines show the standard deviations.

The dotted lines magnetic field intensity.

Correlation coefficients obtained comparing the longitudinal profiles of monthly lower stratosphere brightness temperature climatology and the magnetic field intensity near surface.



Ionizaion rates as a function of altitude



Vampola, A., and D. Gorney (1983), Electron Energy Deposition in the Middle Atmosphere, J. Geophys. Res., 88(A8), 6267-6274.

Fig. 8. Plots of ionization rates as a function of altitude for the average electron precipitation spectra shown in Figures 1 and 2. The peak around 80 km is due to the primary electrons, whereas the secondary peak around 40 km is due to bremsstrahlung X rays.

7/9/2008

SABER/TIMED Data Vertical O3 and NO Profiles





Figure:

The left panels show the vertical profiles inside and outside the magnetic anomaly of the O3 and NO volumetric emission rate, and, temperature for the latitude 350 S to 300 S.

The right panels show the difference between the profiles inside and outside the magnetic anomaly.

Numerical Experiment

Experimental Configuration

- We use the Goddard Institute for Space Studies GCM Model E version
 - run at 4x5 degree horizontal resolution with 20 vertical layers extending up into the stratosphere.
 - GISS Dynamic ocean model (Russell ocean model). Resolution (4x5xL13)
- The basic physics of the model are similar to those used in recent climate studies.
- **Dynamics:**

The scheme is leap frog in time with an initial 2/3 time step every 8 leap-frog steps to prevent solution splitting. T

The dynamics are based on the dry physics (no water vapour effects in the pressure gradient calculation) and use potential temperature as the advected variable.

Two experiments:

Realistic three-dimensional distribution of ozone

Zonal mean ozone (no effects of the zonal assimetry)

- The atmospheric model is coupled to the GISS Dynamic ocean model (Russell). 13 levels.
- 62 years (the initial 48 years of each experiment, during which the model was reaching equilibrium, were discarded).

Schmidt, G.A., R. Ruedy, J.E. Hansen, I. Aleinov, N. Bell, M. Bauer, S. Bauer, B. Cairns, V. Canuto, Y. Cheng, A. Del Genio, G. Faluvegi, A.D. Friend, T.M. Hall, Y. Hu, M. Kelley, N.Y. Kiang, D. Koch, A.A. Lacis, J. Lerner, K.K. Lo, R.L. Miller, L. Nazarenko, V. Oinas, Ja. Perlwitz, Ju. Perlwitz, D. Rind, A. Romanou, G.L. Russell, Mki. Sato, D.T. Shindell, P.H. Stone, S. Sun, N. Tausnev, D. Thresher, and M.-S. Yao, 2006: Present day atmospheric simulations using GISS ModelE: Comparison to in-situ, satellite and reanalysis data. J. Climate, 19, 153-192, doi:10.1175/JCLI3612.1.





Stratospheric Ozone



The boundary between stratospheric and tropospheric ozone: 125-hPa level.

Annual mean Sea Surface Temperature

Difference between symmetric and asymmetric configurations

SEA SURFACE TEMPERATURE



La Nina Signature



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South Atlantic Convergence Zone



Surface Temperature

Difference between symmetric and asymmetric configurations

GISS Surface Temperature Analysis







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ENSO evolution from the Late Glacial through the Holocene



Figure 5. Results of time series analysis (16) of the gray-scale record (Fig. 3C); age ranges are in calendar years before the present. Because of breaks in the core (Fig. 3A), we performed the time series analysis on individual sections of the core and treated these as discrete time series. There is a clear spectral evolution from the Late Glacial through the Holocene with the ENSO band (19), which progressively achieves its modern strength by ~5000 cal yr B.P.

RODBELL et al. (1999), Science, Vol. 283. no. 5401, pp. 516 - 520 DOI: 10.1126/science.283.5401.516