Understanding Solar Flare Heating: Week Four Update

BROCK PARKER

DR. CHUNMING ZHU

Background

- Solar Flares
 - Violent release of radiation
 - Related to magnetic field

- Energy
 - Magnetic field reconnection
 - OProportional to footprint

Goals

- Model Synthetic Flares
 - Constrain heating functions
 - •Flare loop generation
 - Synthetic light curves
 - One flare

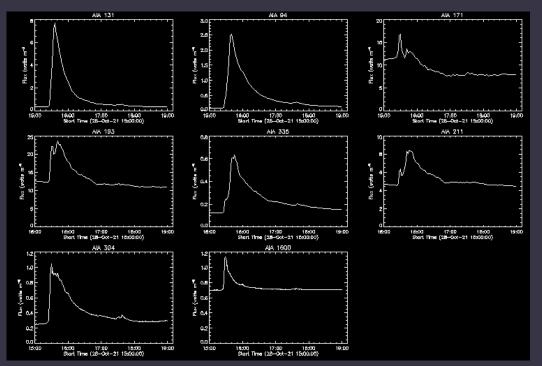
- Synthetic Images
 - ○0D and 3D combination
 - Formation insight

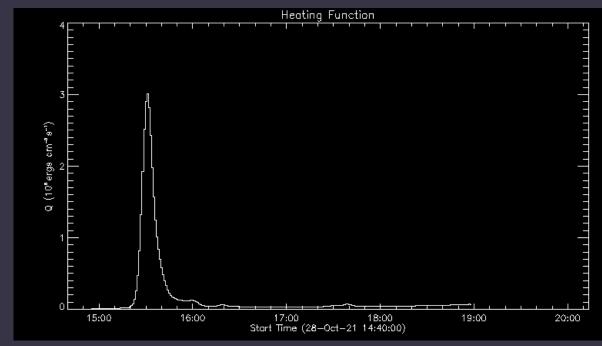
Process

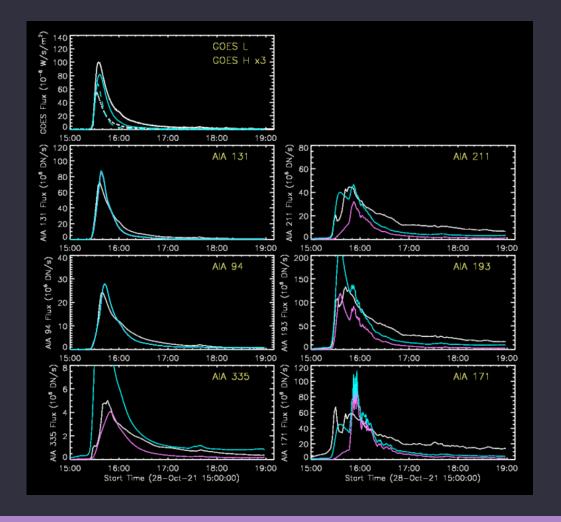
- Download and align data
- Set parameters
- Determine bright pixels
- Generate lightcurves for each separate filter
- Calculate loop lengths
- Calculate the heating function for each

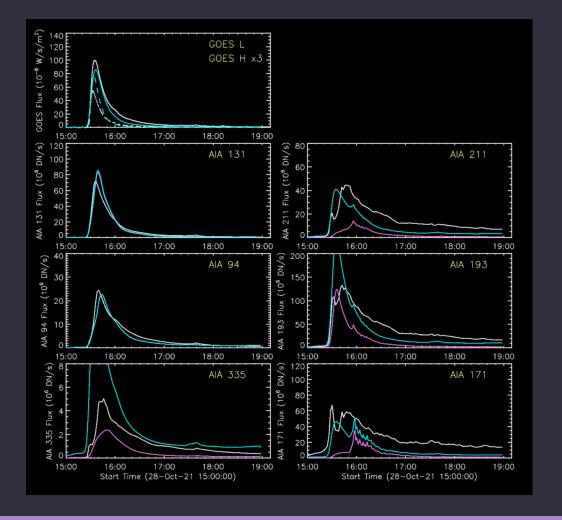
- Run EBTEL time evolution of density,pressure, and temperature for eachloop
- Generate the simulated lightcurves
- Compare to observations
- Combine with EBTEL
- Generate '3d' image of full time evolved loops

Previously Light Curves Heating Functions

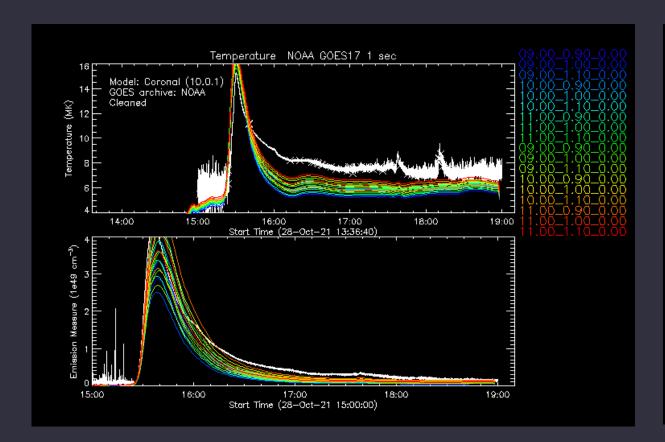


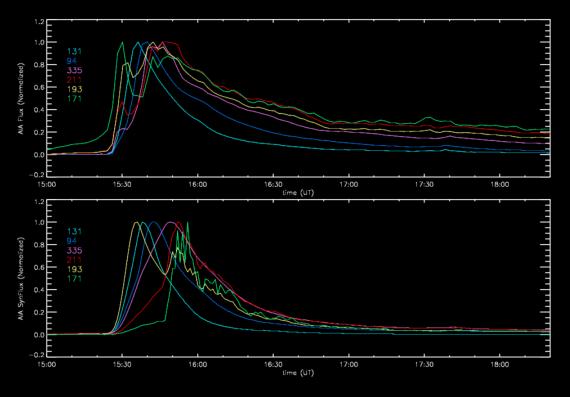




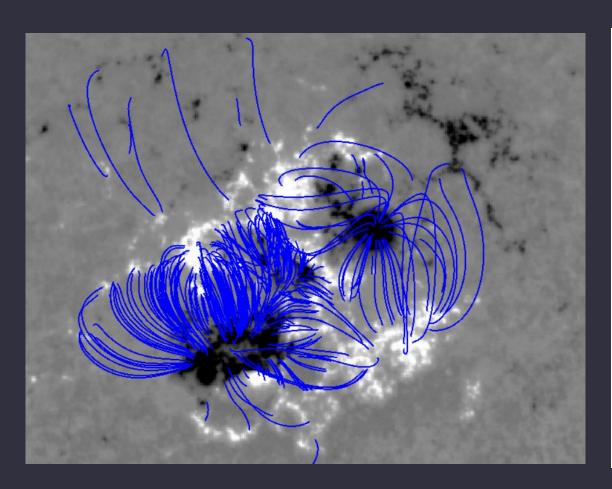


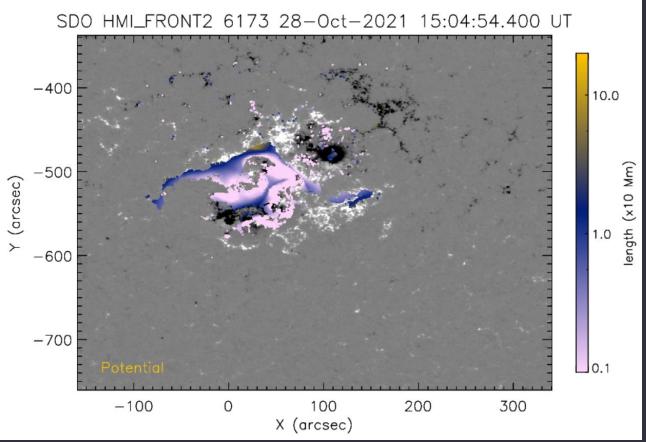
Progress EBTEL





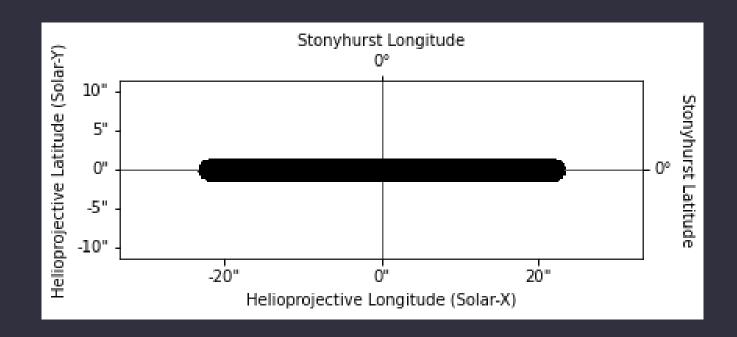
Progress EBTEL

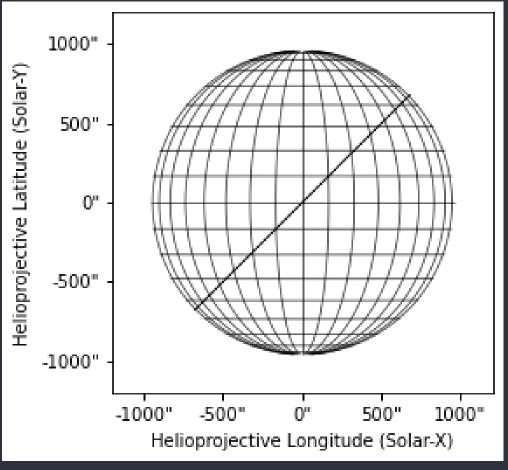


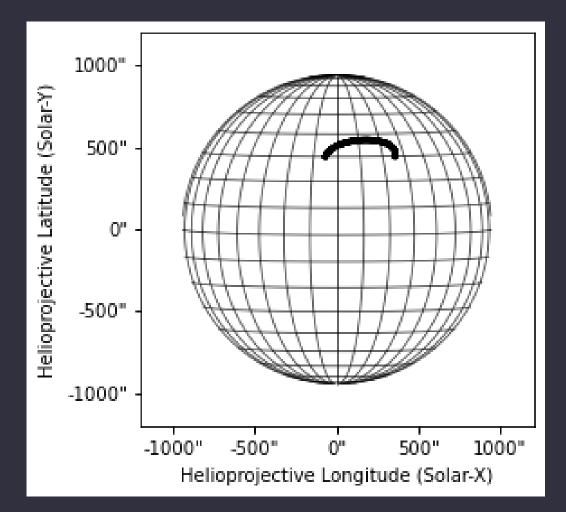


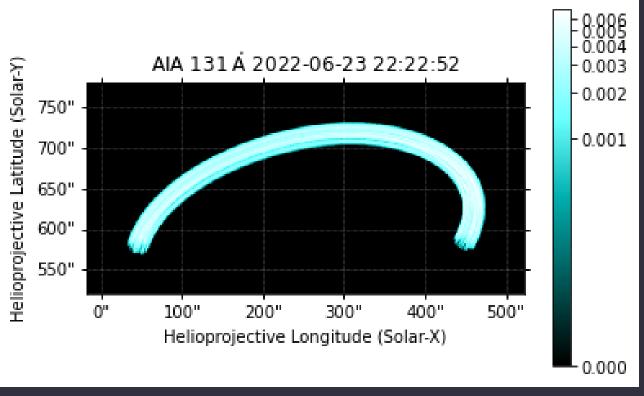
Process

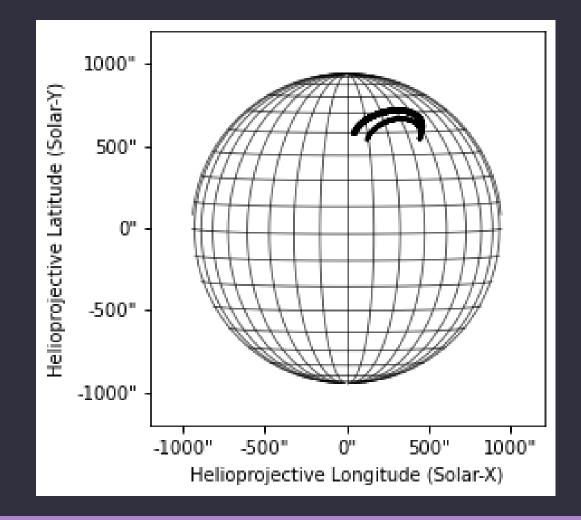
Extrapolation

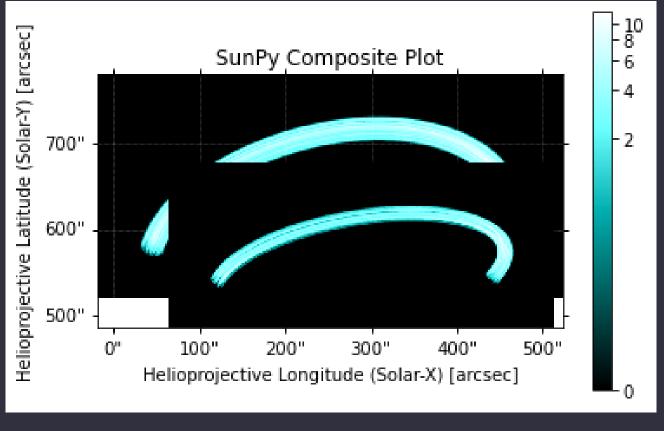


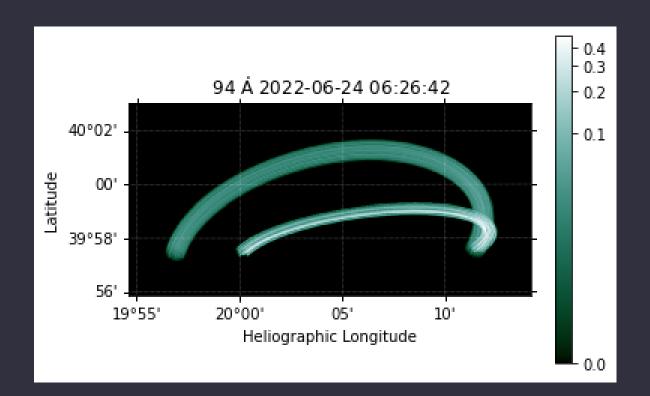


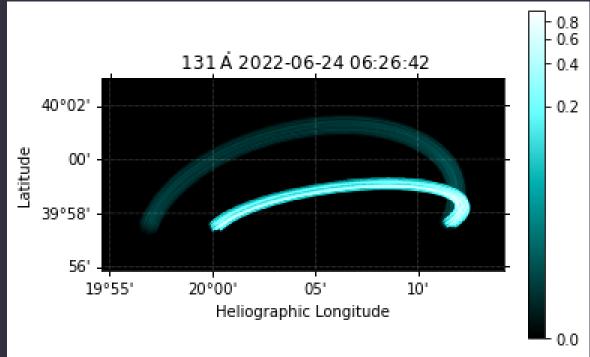


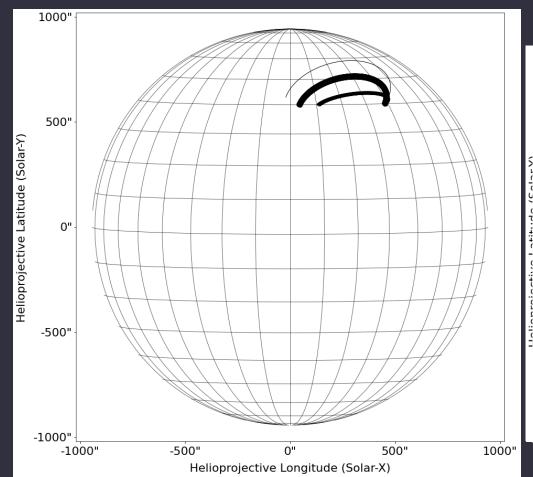


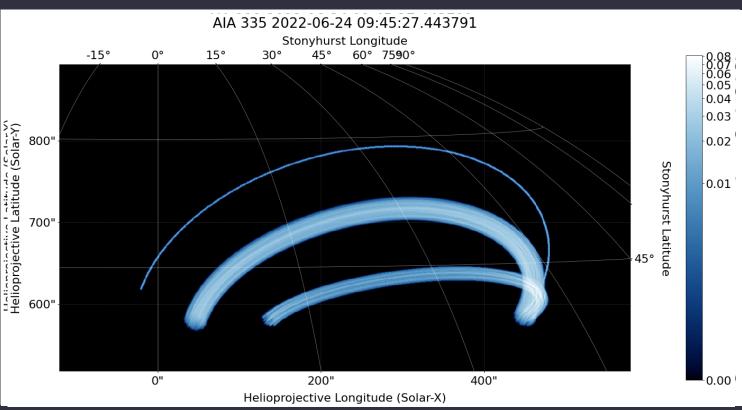












References

- oJ. A. Klimchuk, et al. 2008, ApJ, **682**, 1351.
- oJiong Qui, et al. 2012, ApJ, **752**, 124.
- oChunming Zhu, et al. 2018, ApJ, **856**, 27.
- oW. T. Barnes, et al. 2019, ApJ, 880, 56.